

Van Allen Probes: Storms, Substorms, and Radiation Belt Structure

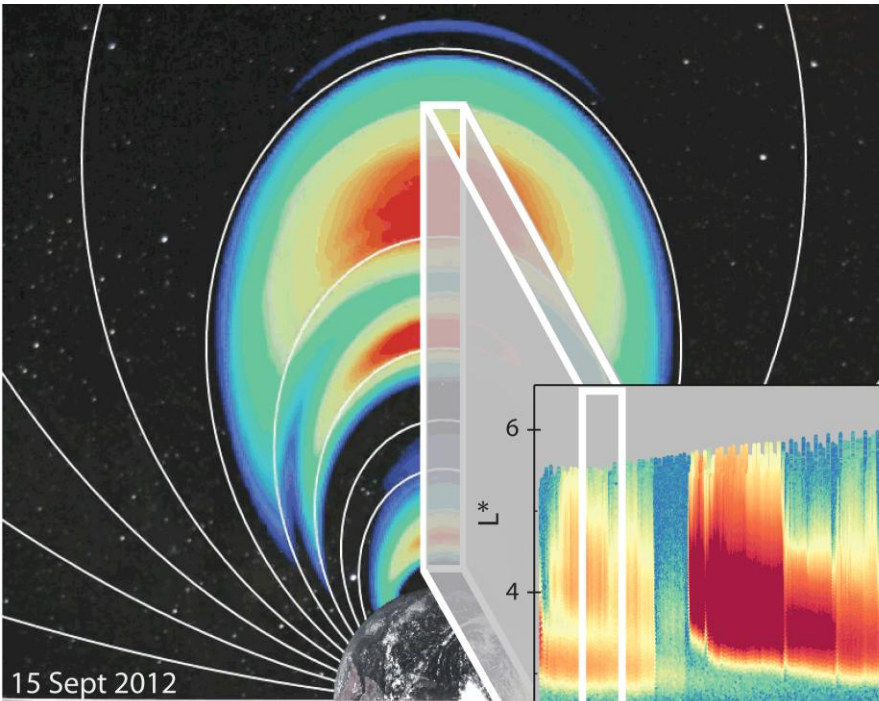
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Department of Physics
University of Colorado - Boulder

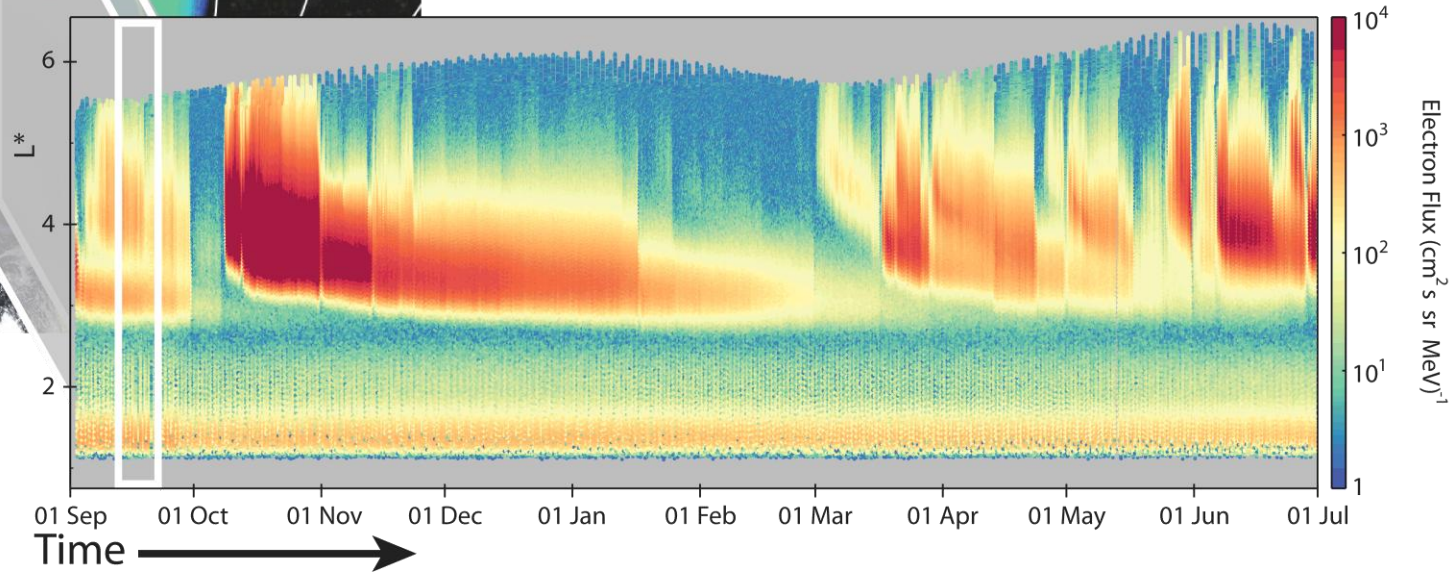
Contributors: A. Jaynes, V. Hoxie, X. Li, S. Kanekal, R. Thorne, J. Foster, P. Erickson,
D. Malaspina, J. Wygant, W. Kurth, W. Li, Q. Ma, L. Blum, Q. Schiller, A. Gerrard, L. Lanzerotti

REPT Observations:

Acceleration, Remanence, and Sudden Loss



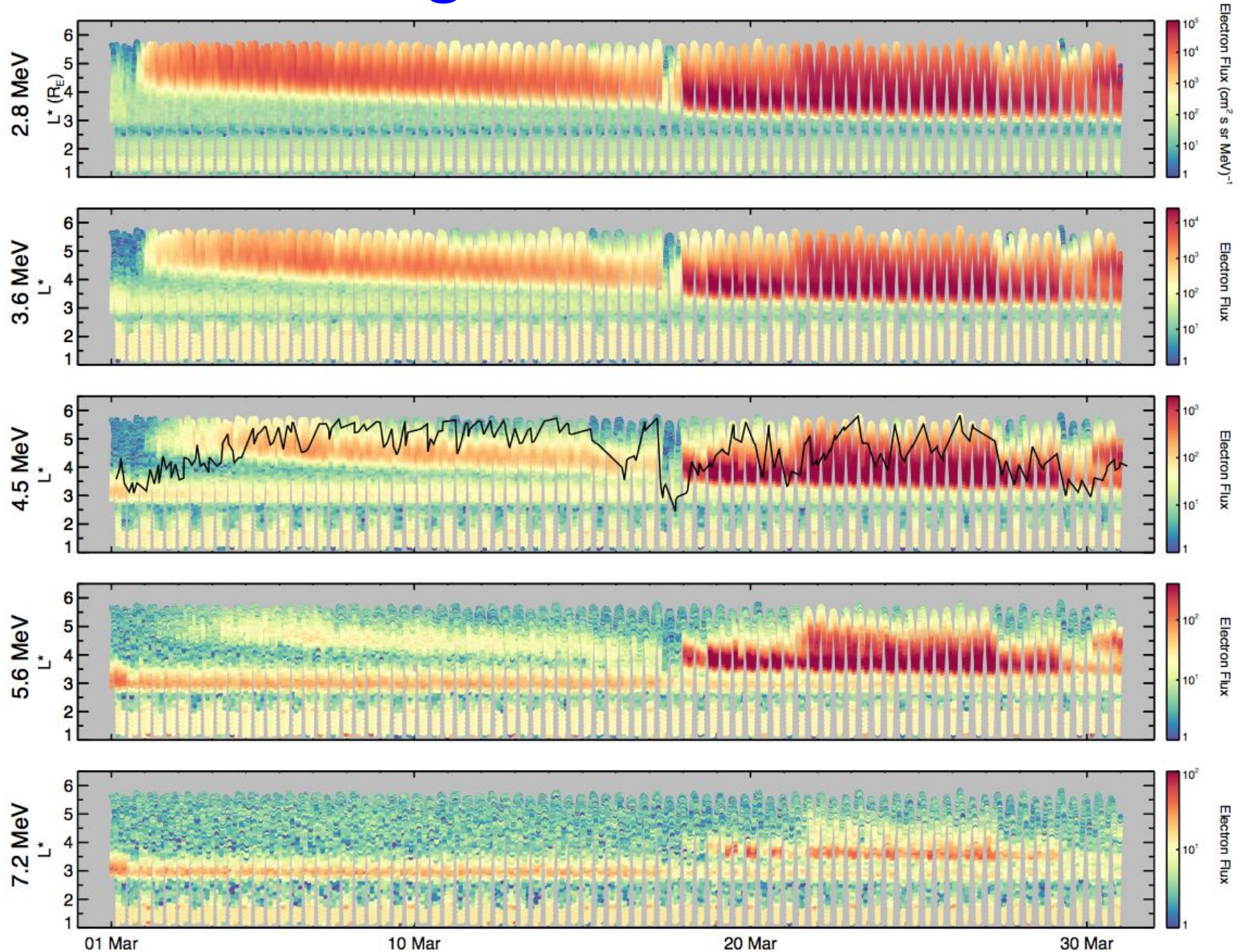
4.5 MeV electron fluxes



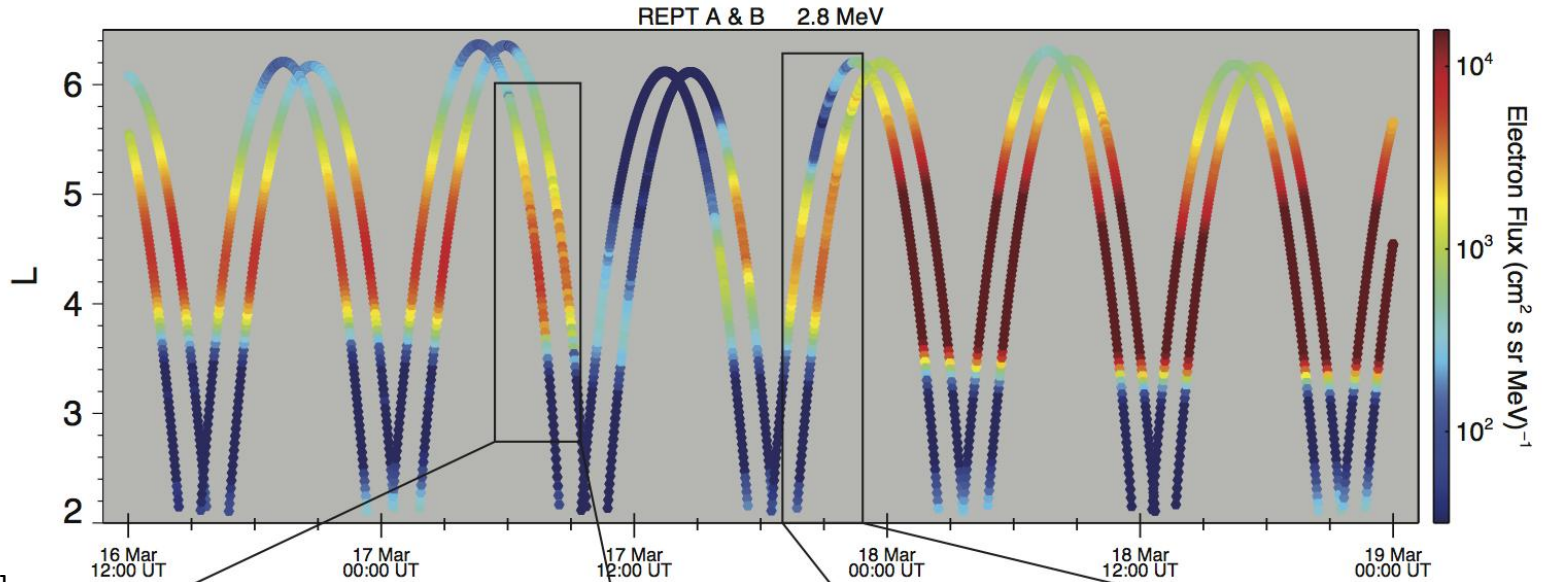
Baker et al., 2013



Fascinating Period: March 2013

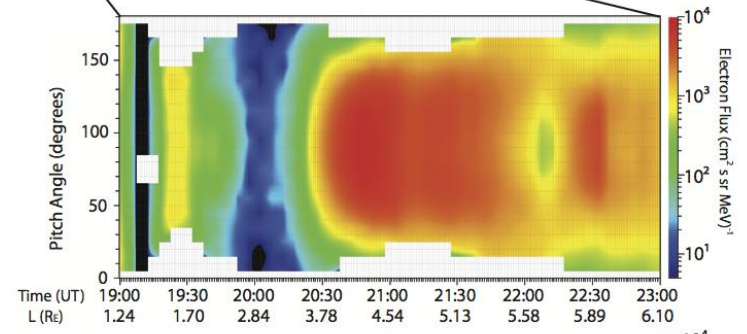
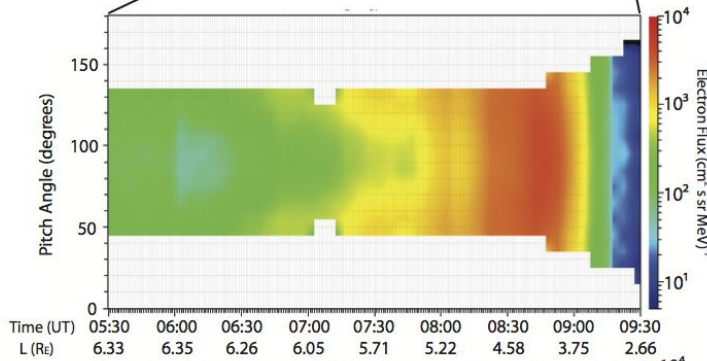


Exquisite View of Flux Dropout and Recovery

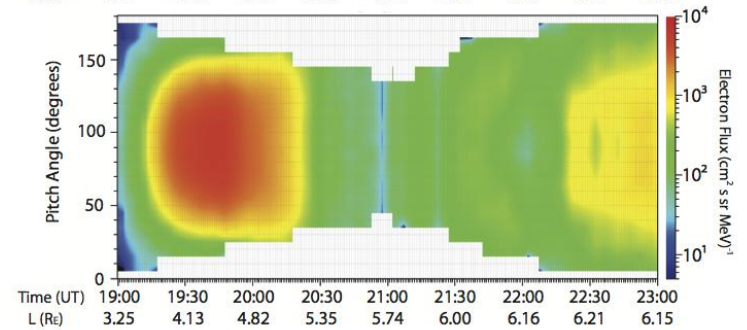
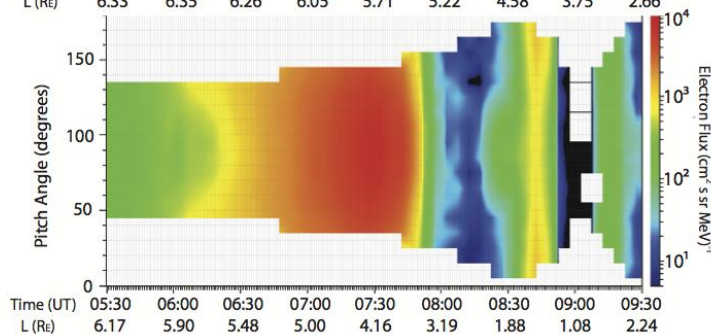


[Baker et al., GRL,2014]

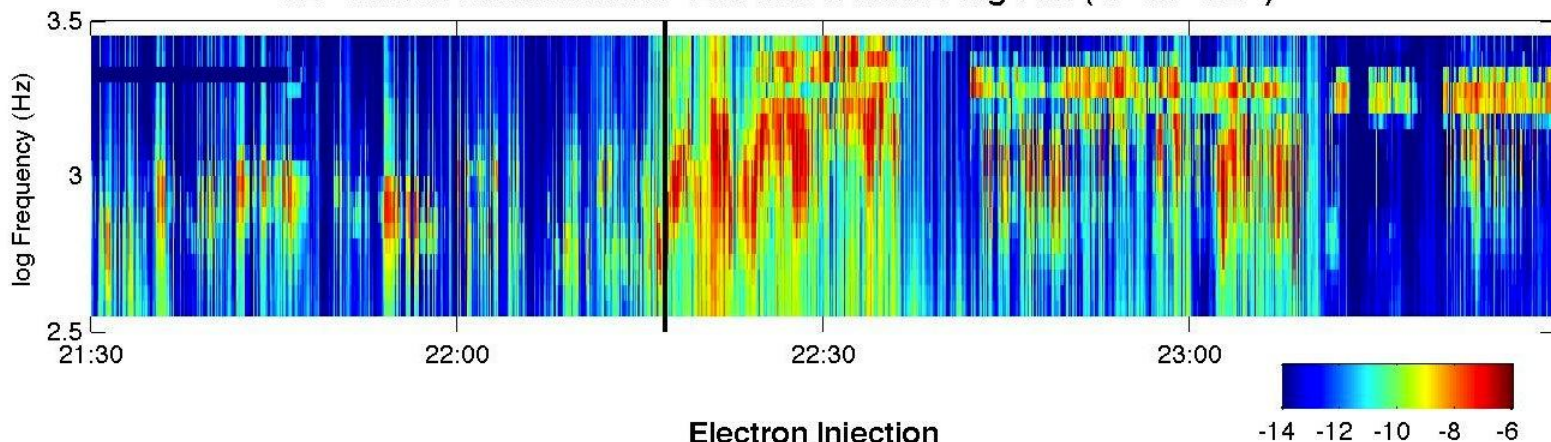
REPT A
PAD



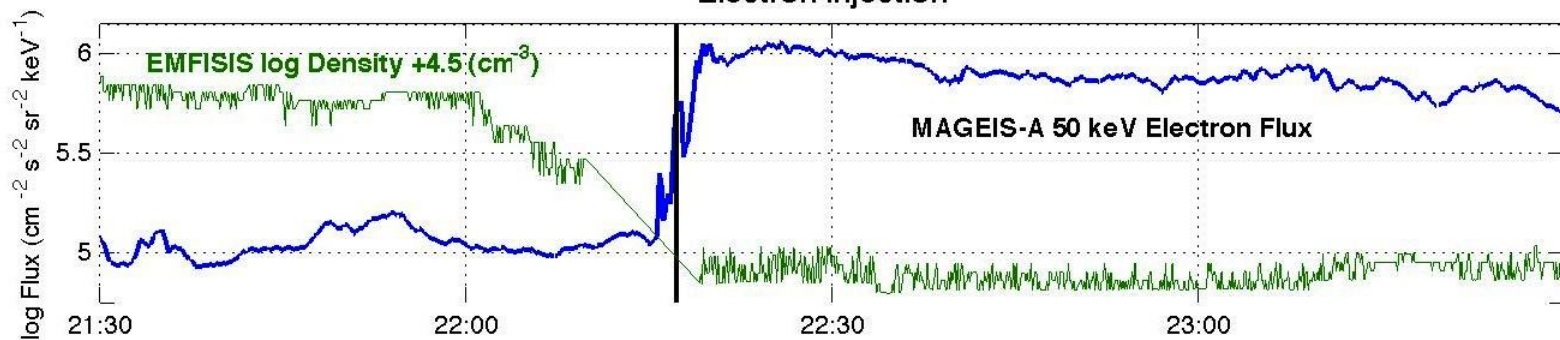
REPT B
PAD



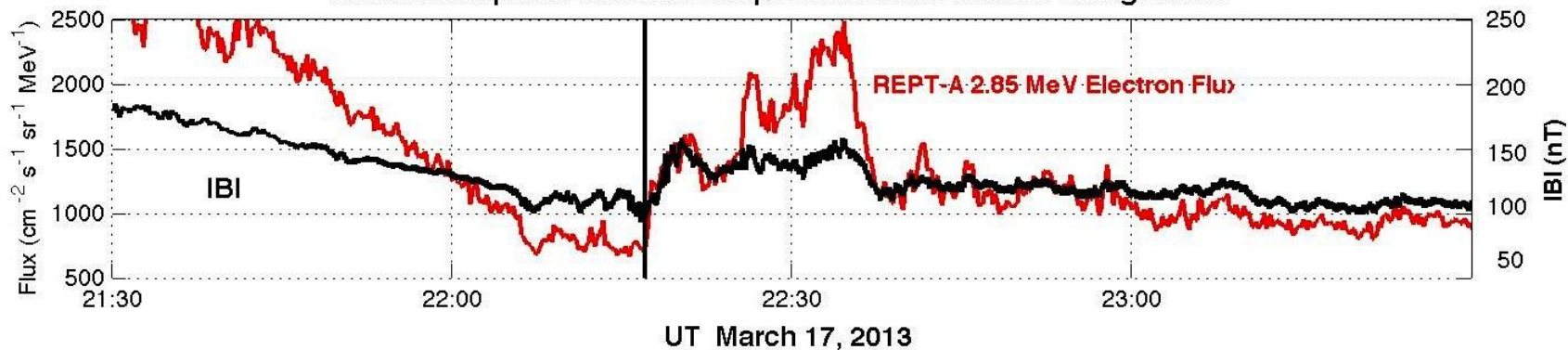
VLF Chorus Enhancement: EMFISIS-A total E log PSD ($V^2 m^{-2} Hz^{-1}$)



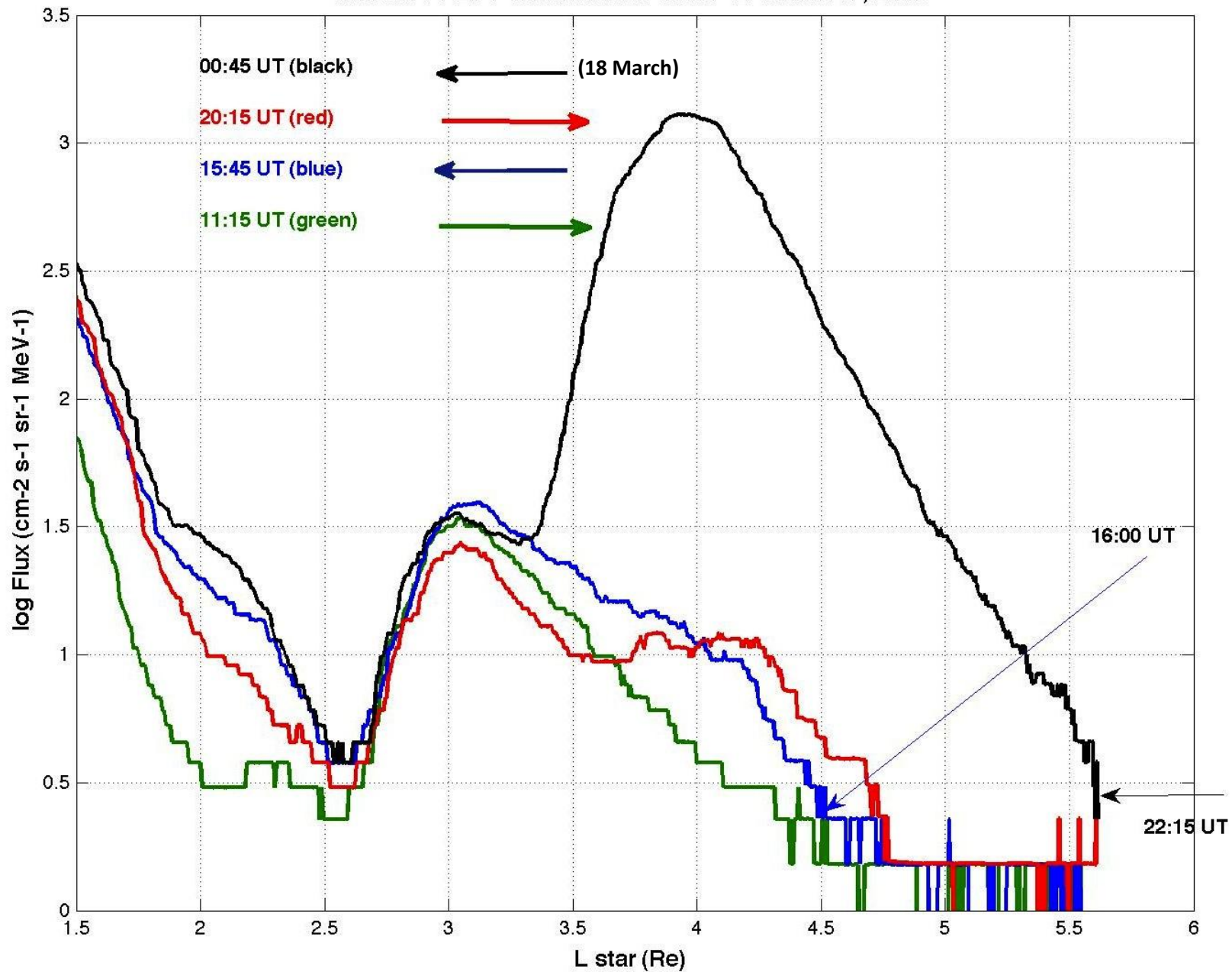
Electron Injection



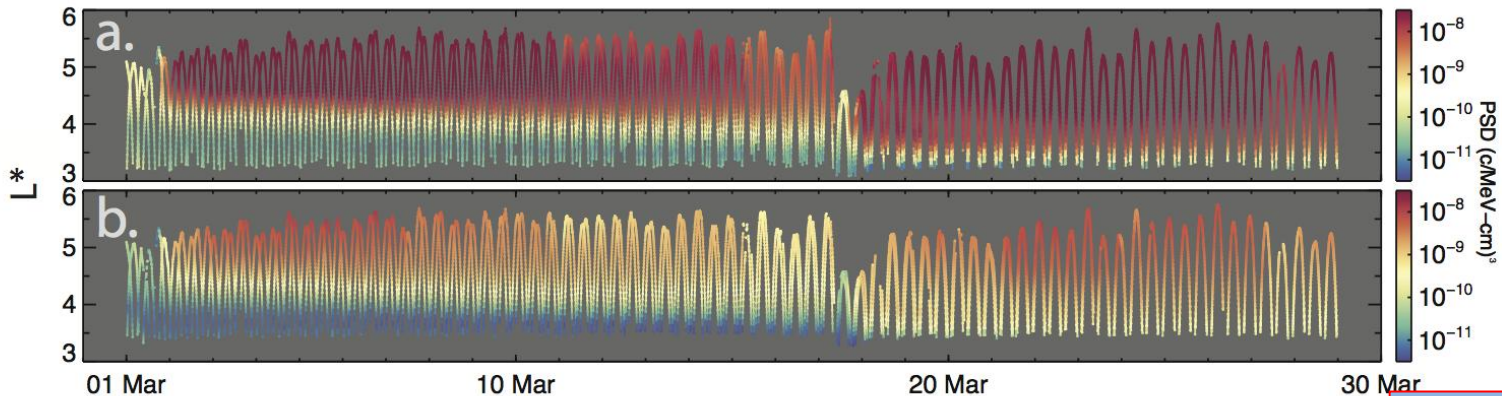
Substorm Dipolarization & Prompt Relativistic Electron Energization



4.50 MeV REPT electron Flux RBSP-B March 17, 2013

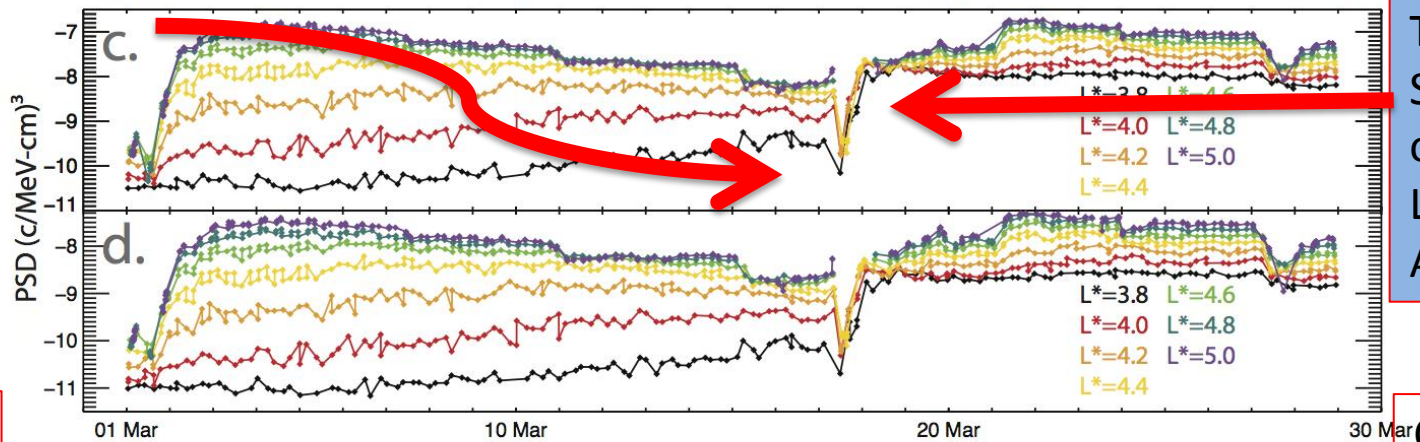


$\mu=2879$
MeV/G



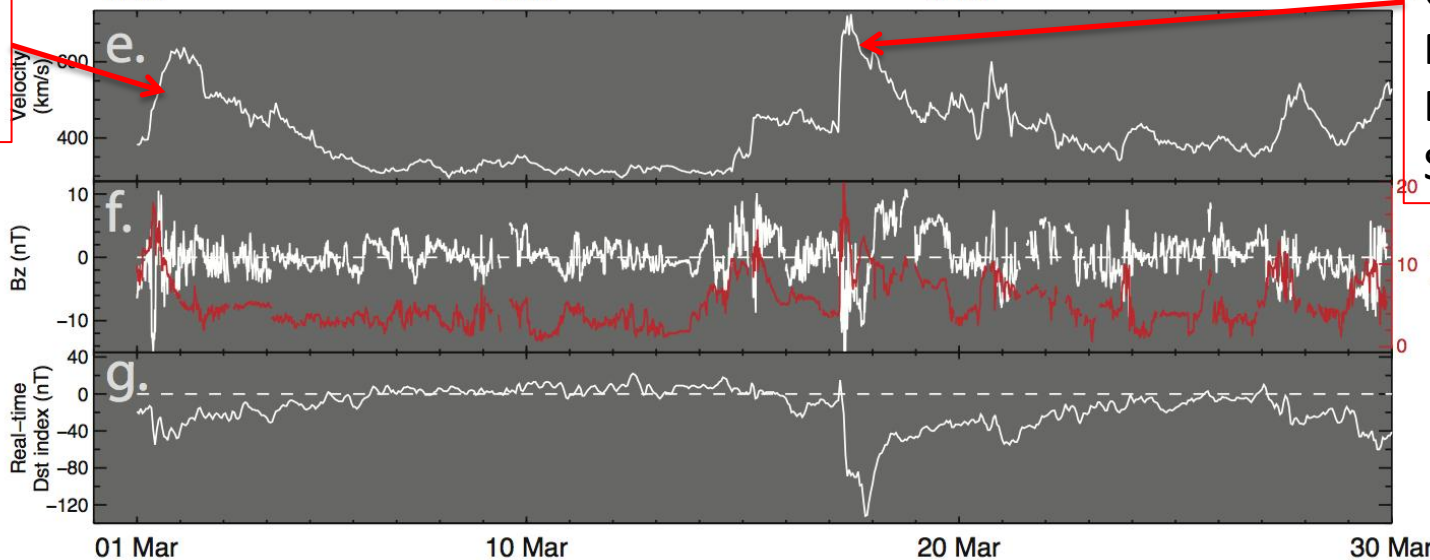
The
Signature
of
Radial
Diffusion

$\mu=3433$
MeV/G

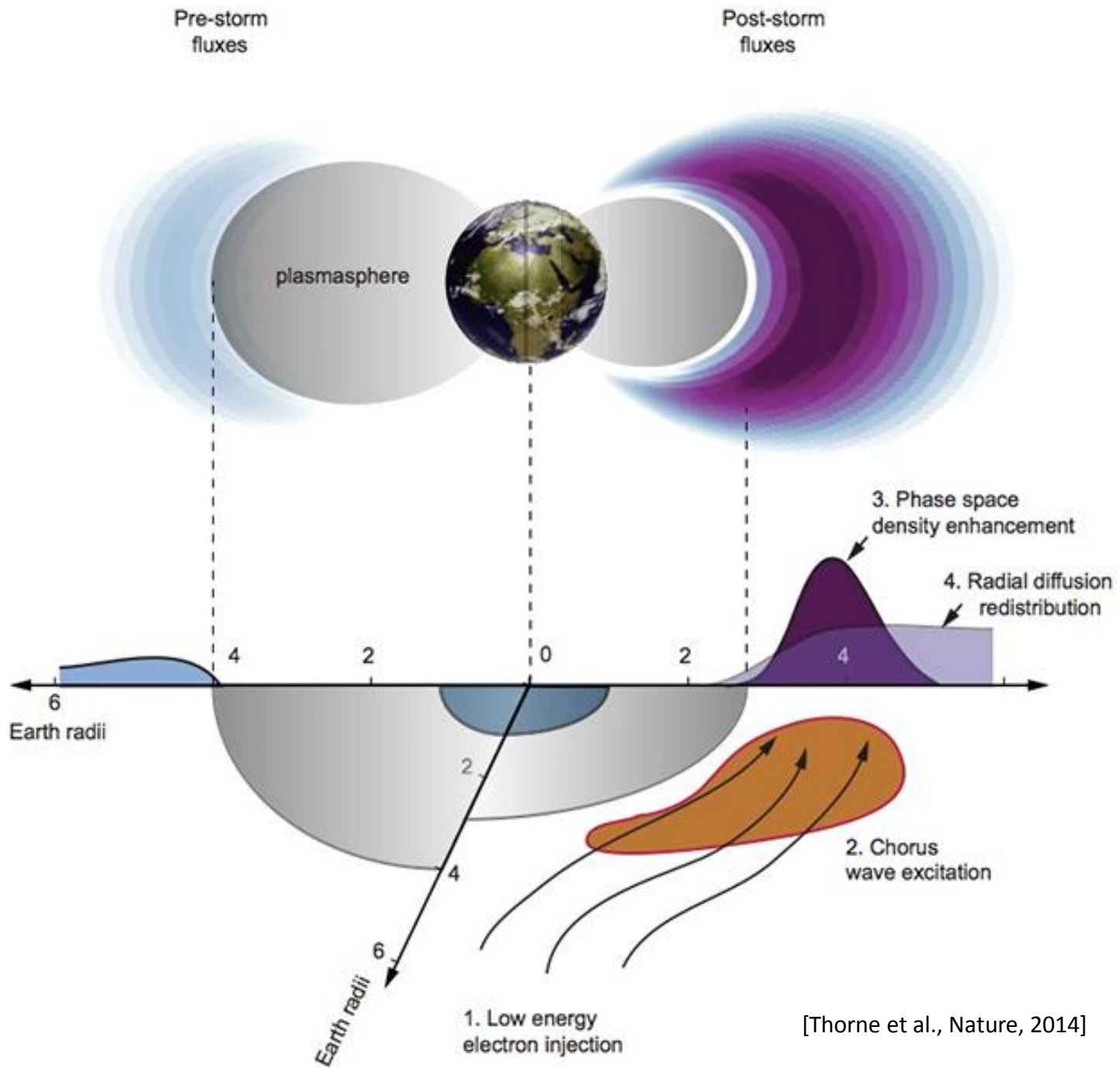


The
Signature
of
Local
Acceler.

High-speed
Solar
Wind
Stream

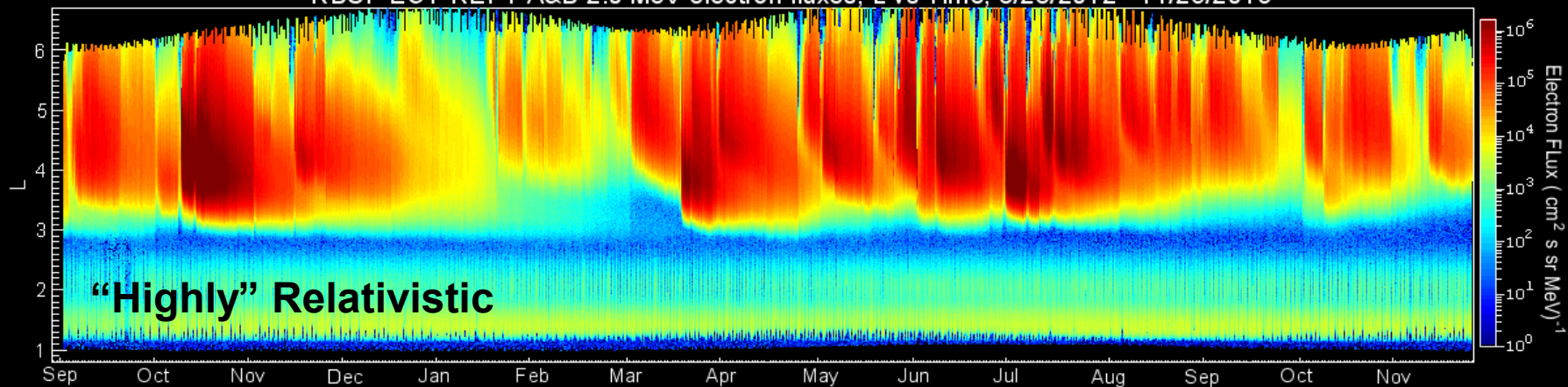


Coronal
Mass
Ejection/
Shock

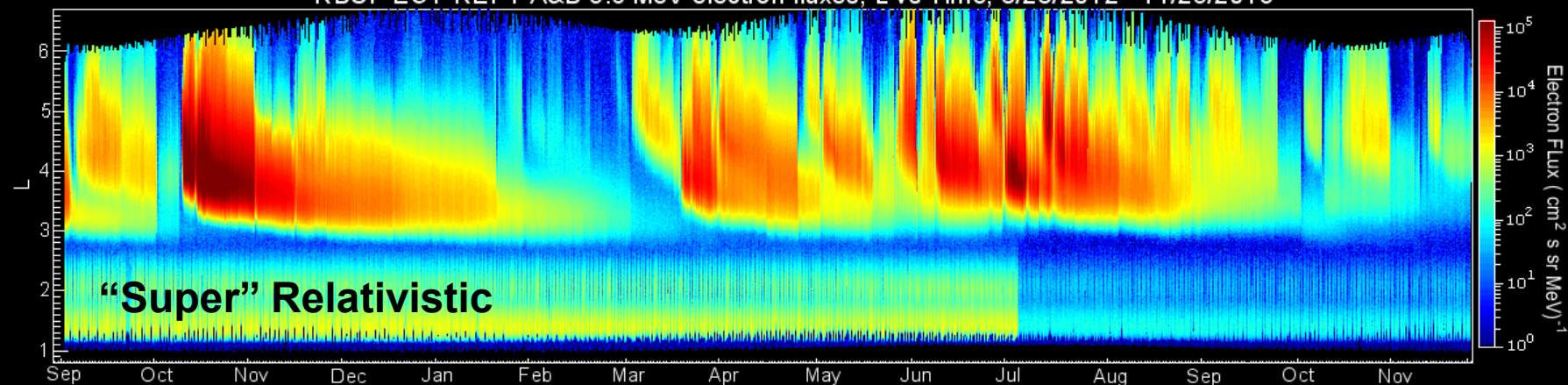


[Thorne et al., Nature, 2014]

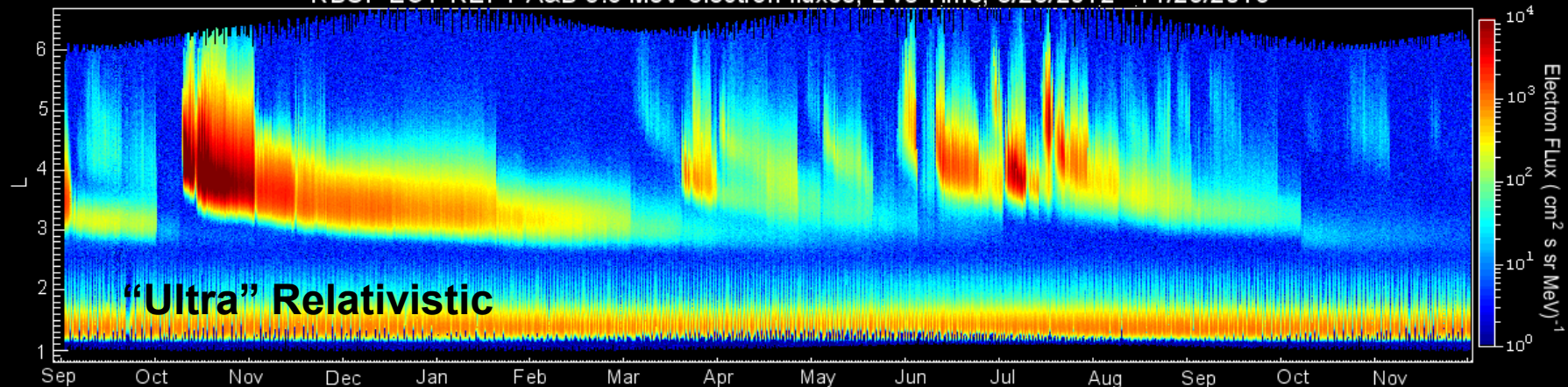
RBSP ECT-REPT A&B 2.0 MeV electron fluxes, L vs Time, 8/28/2012 - 11/25/2013



RBSP ECT-REPT A&B 3.6 MeV electron fluxes, L vs Time, 8/28/2012 - 11/25/2013



RBSP ECT-REPT A&B 5.6 MeV electron fluxes, L vs Time, 8/28/2012 - 11/25/2013



Using full PHA data from REPT

Pulse-height analysis (PHA) packets are recorded onboard at a rate of 83 per second. Much more detail can be determined from this than from the binned flux values calculated onboard.

Method for PHA data analysis (Selesnick, Baker, Jaynes, Li, Kanekal, Hudson & Kress, JGR, 2014):

Valid electron and proton PHA events, that enter through the collimator and stop in detector N , are identified by combined theoretical Landau-Vavilov probability distributions $L(\Delta, E)$ for energy loss Δ given incident energy E :

$$P = \prod_{i=1}^N L(\Delta_i, E_N - E_{i-1})$$

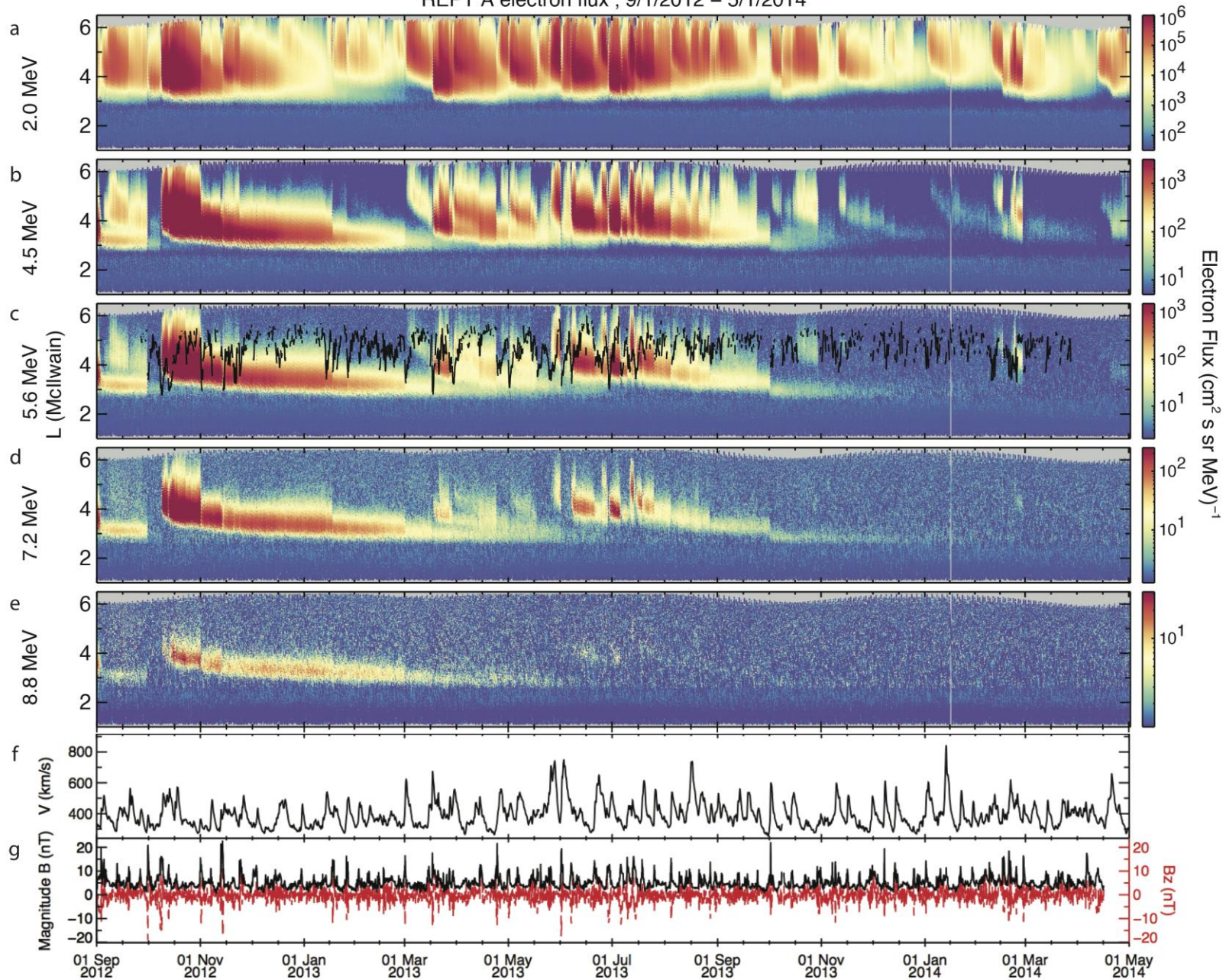
where Δ_i are the measured energy losses in each detector and

$$E_i = \sum_{j=1}^i \Delta_j$$

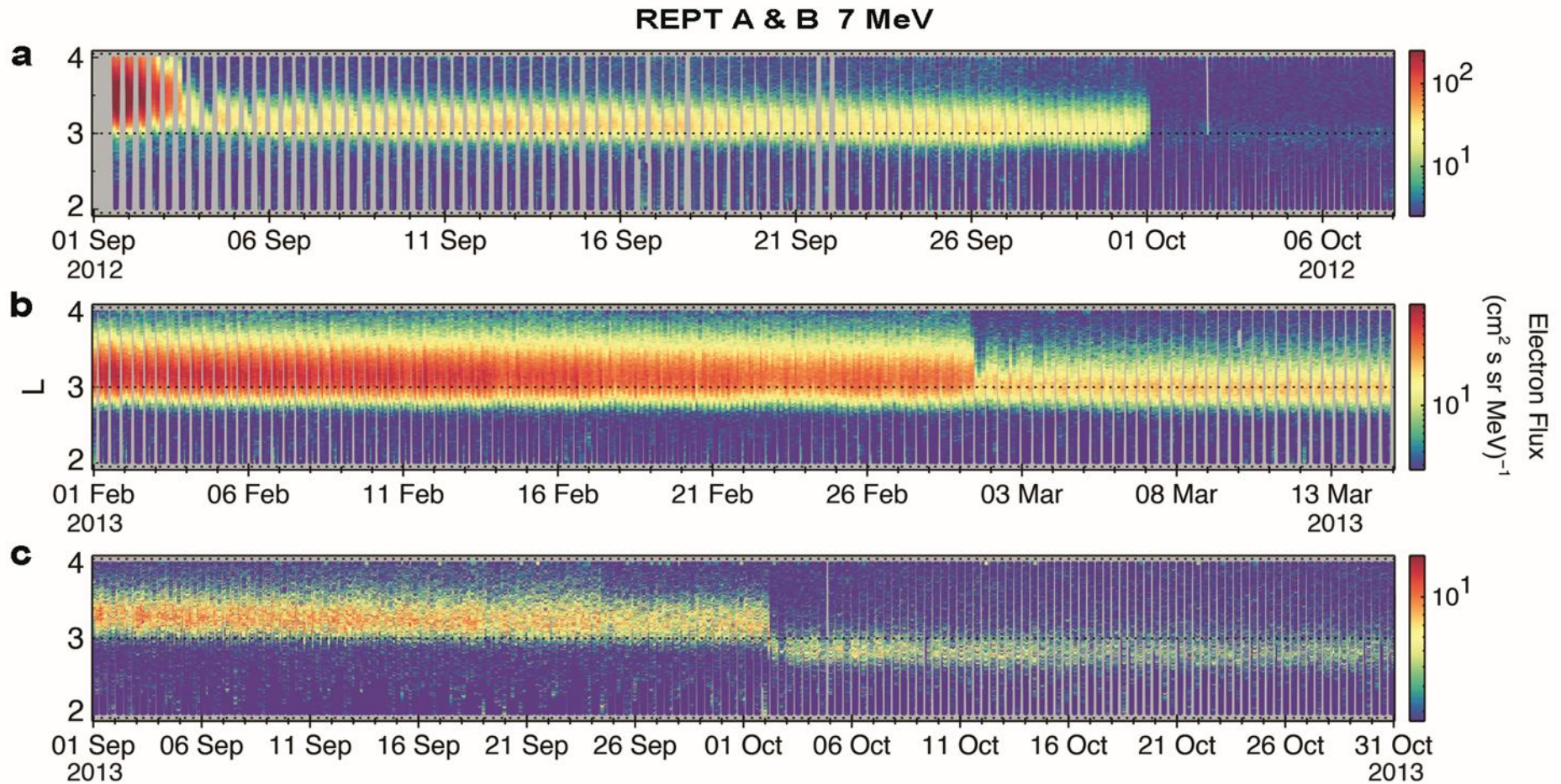
is the total energy loss up to detector i .

Solar Wind Drivers

REPT A electron flux , 9/1/2012 – 5/1/2014

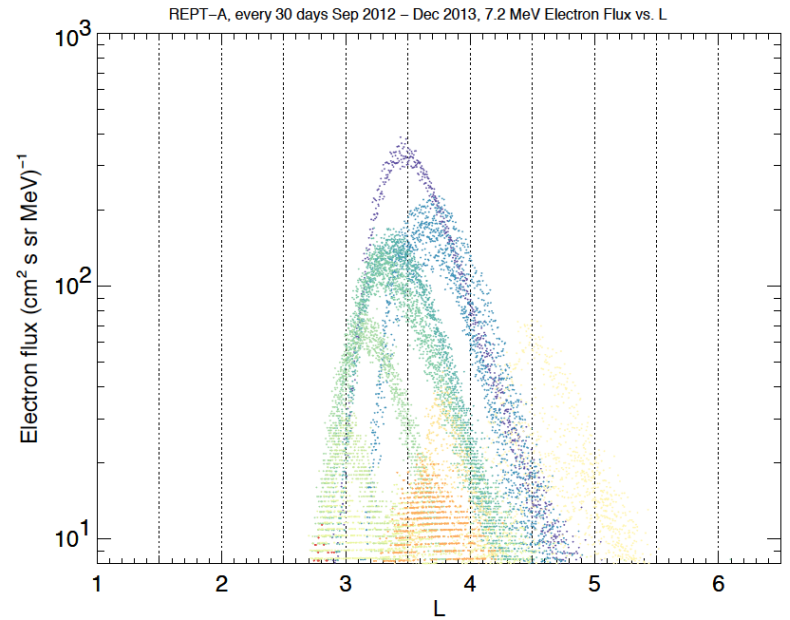
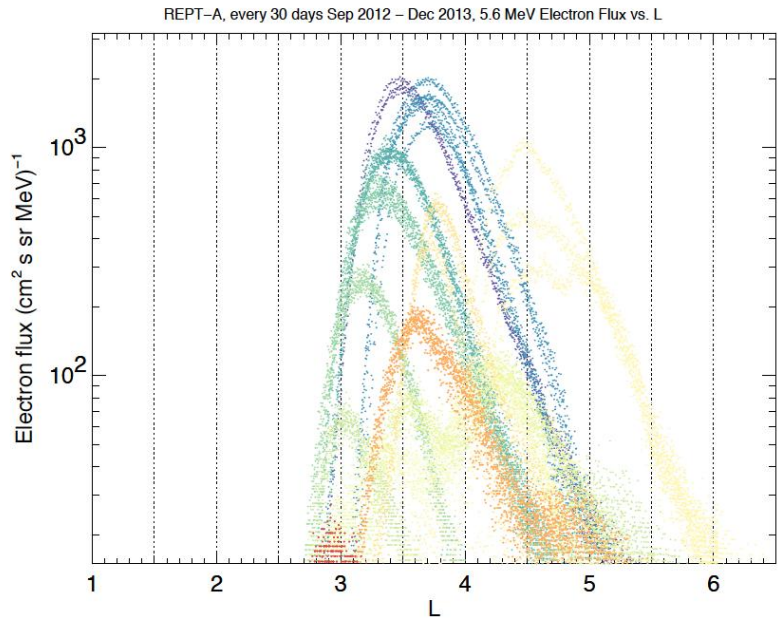
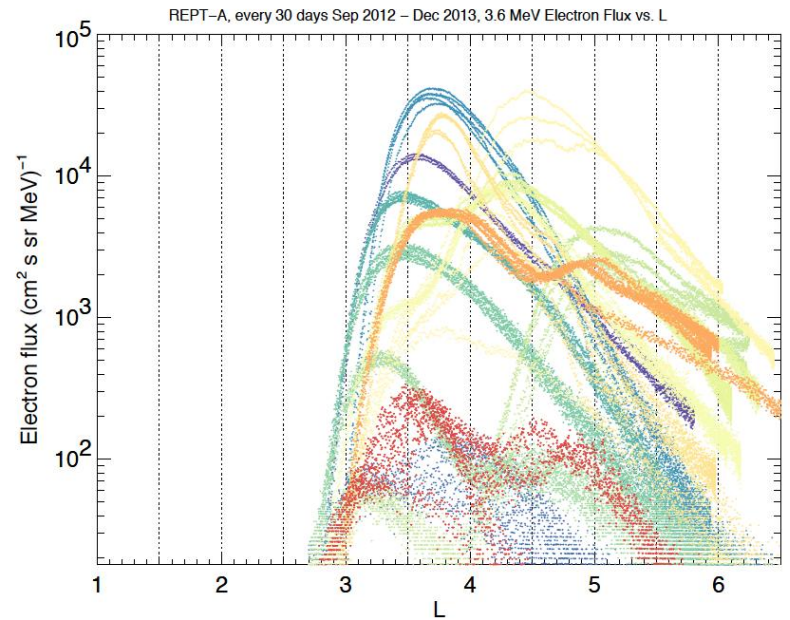
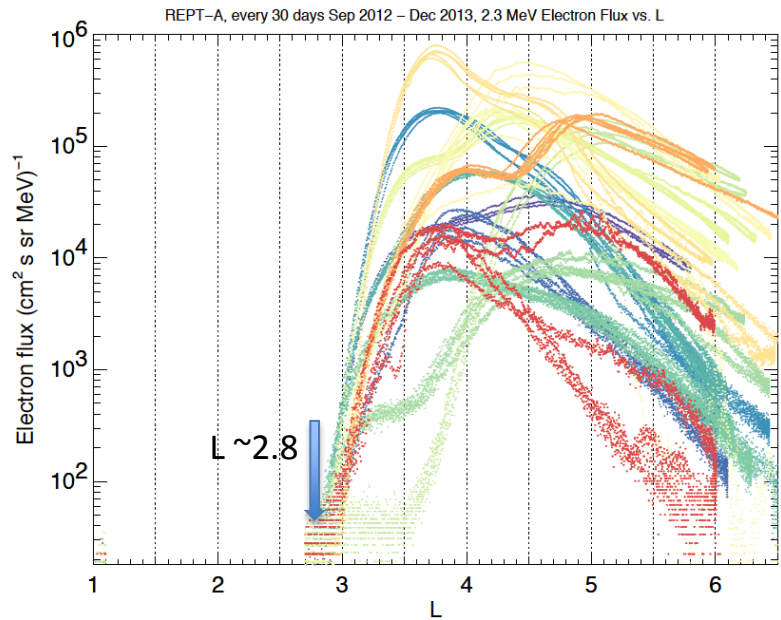


Ultra-Relativistic Electrons: Inner Edge



[Baker et al., Nature, 2014]

Radial profiles of electron flux 2012-2013

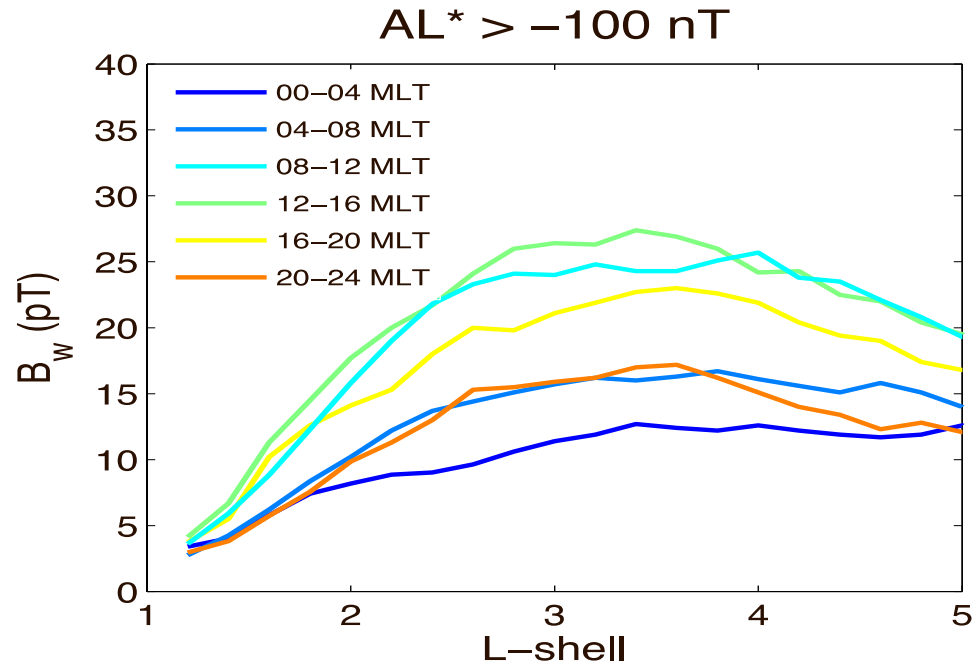


02 Oct 2012 ... color indicates 30 days ...

10 Dec 2013

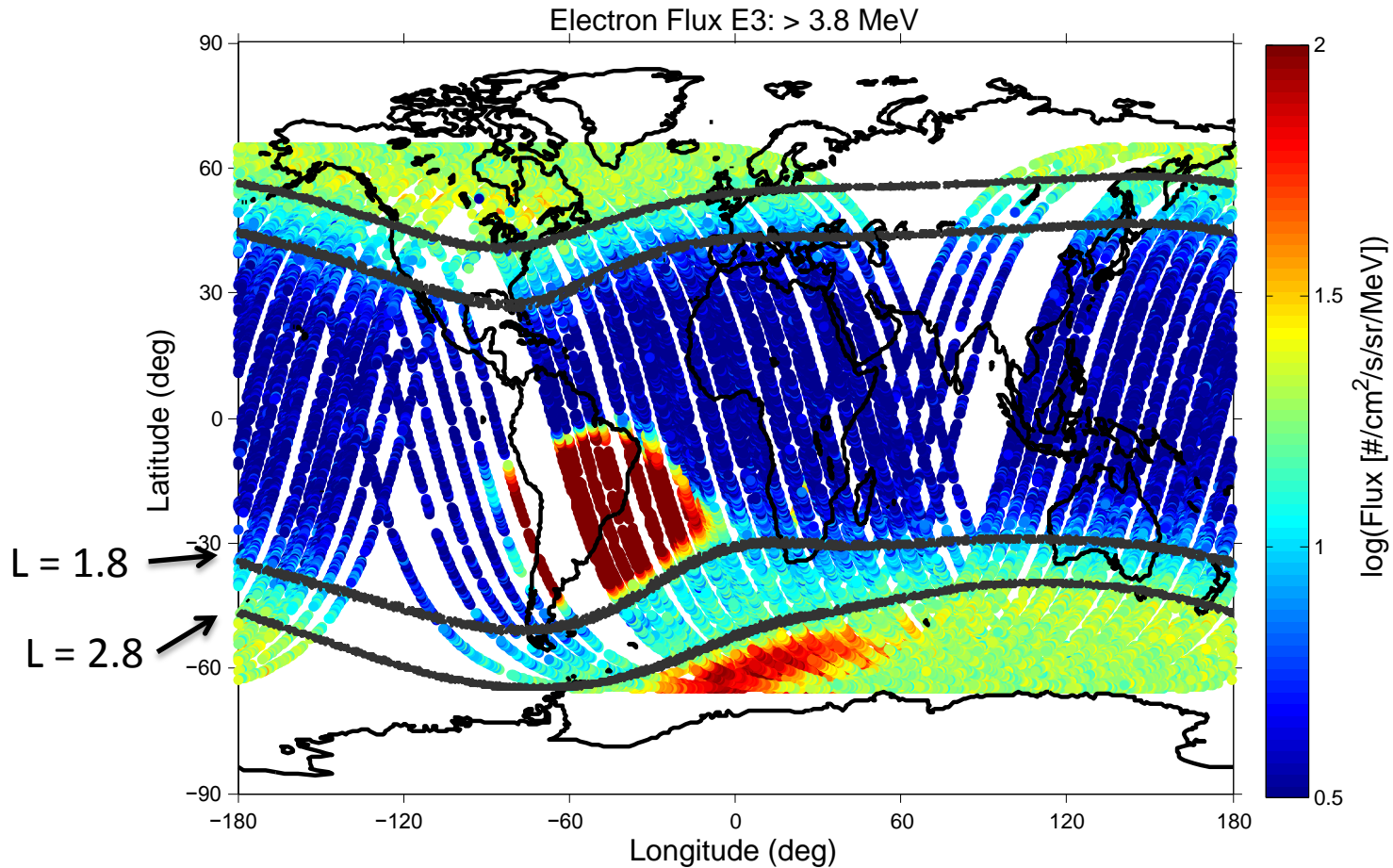
Survey of Plasmaspheric Hiss

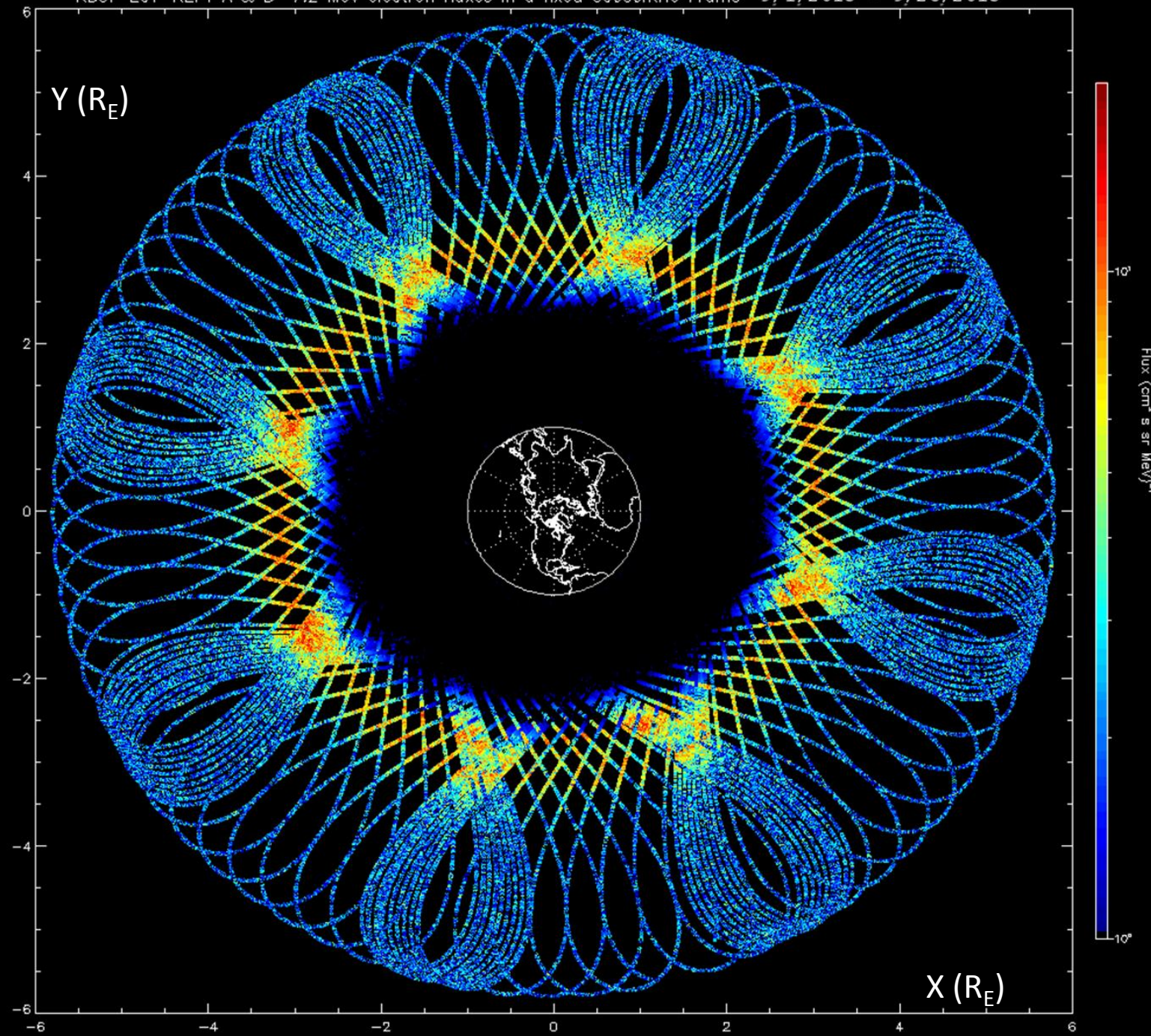
(EMFISIS Data-Wen Li Analysis)

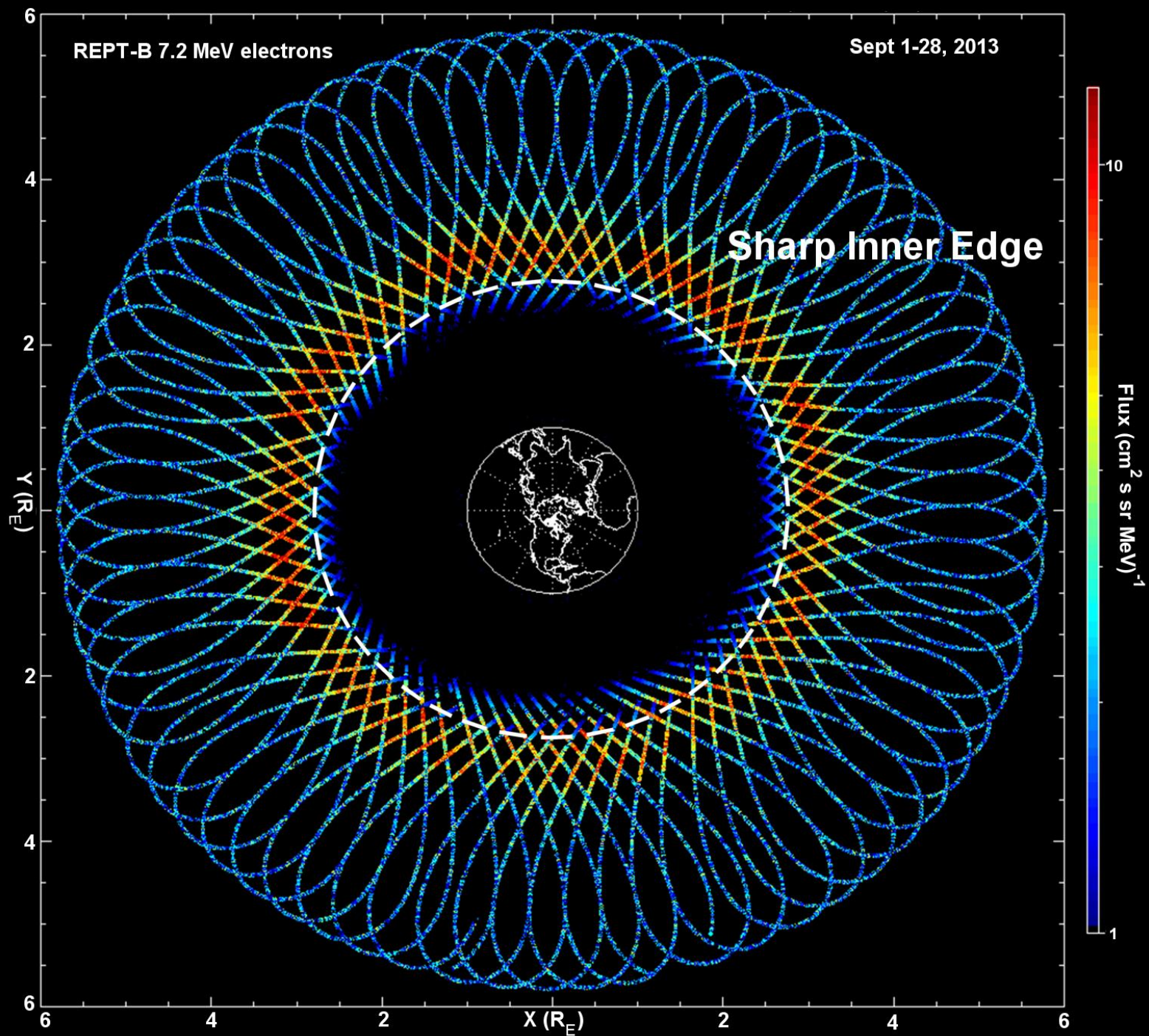


Colorado Student Space Weather Experiment

REPTile Data – September 2013

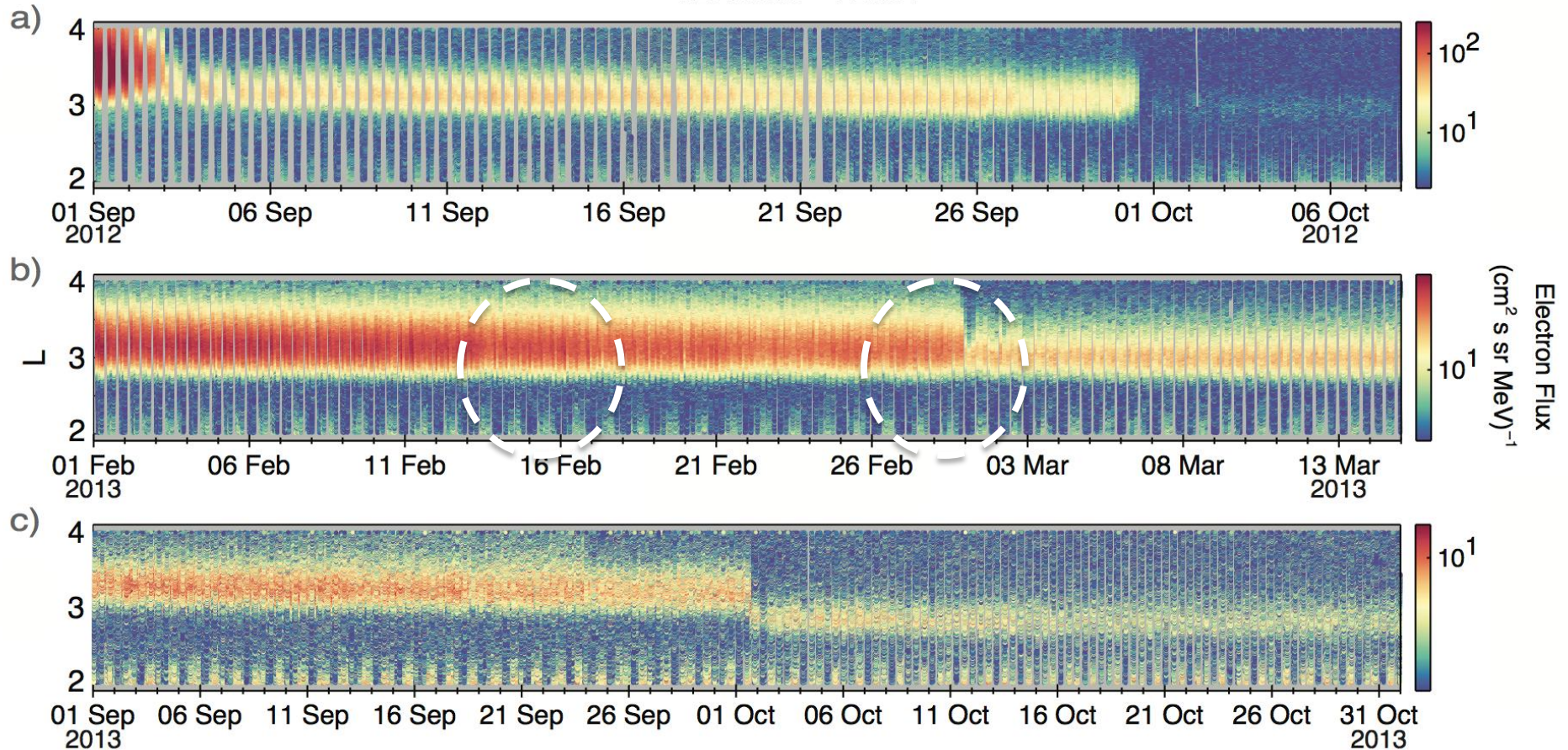


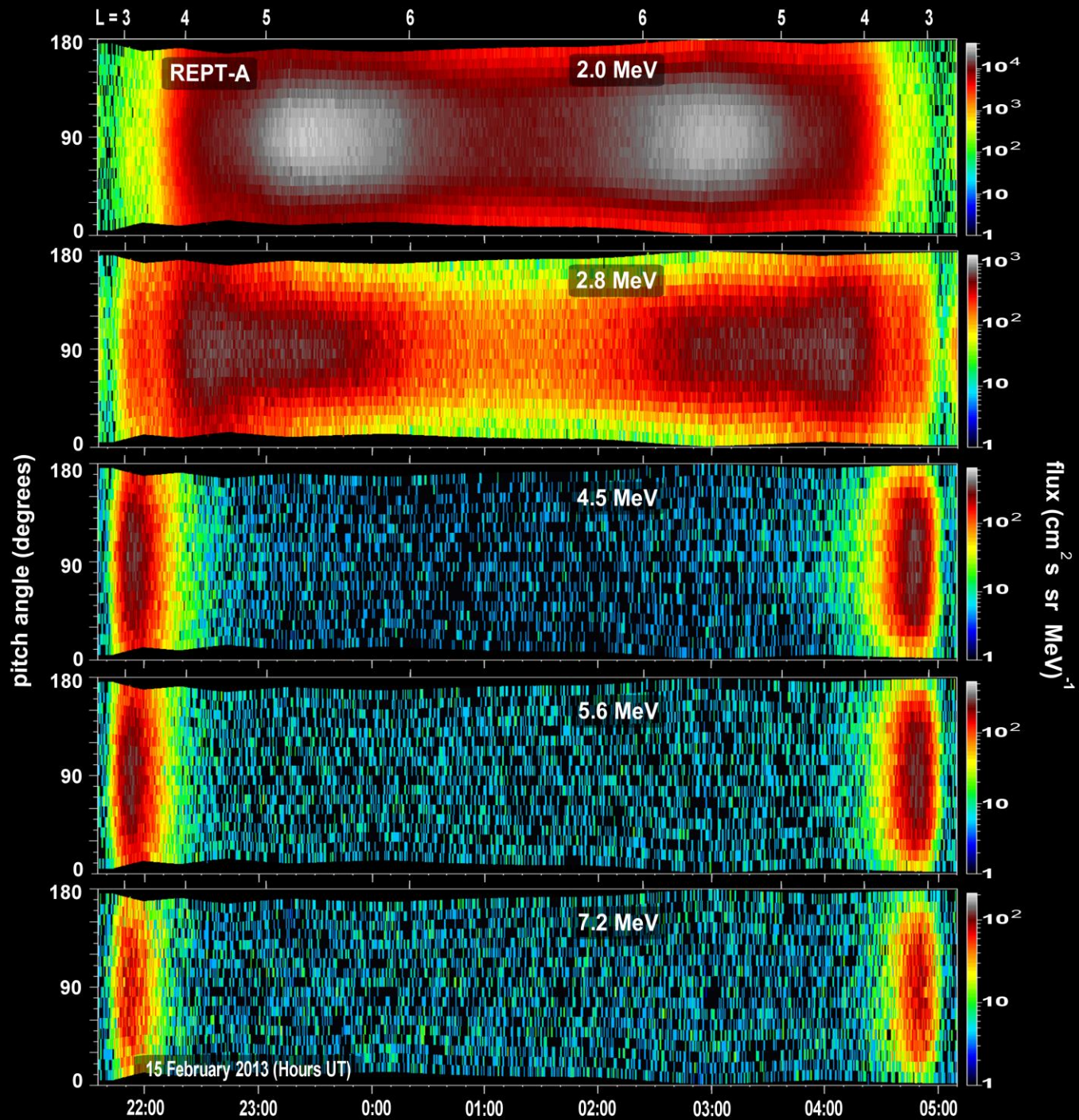


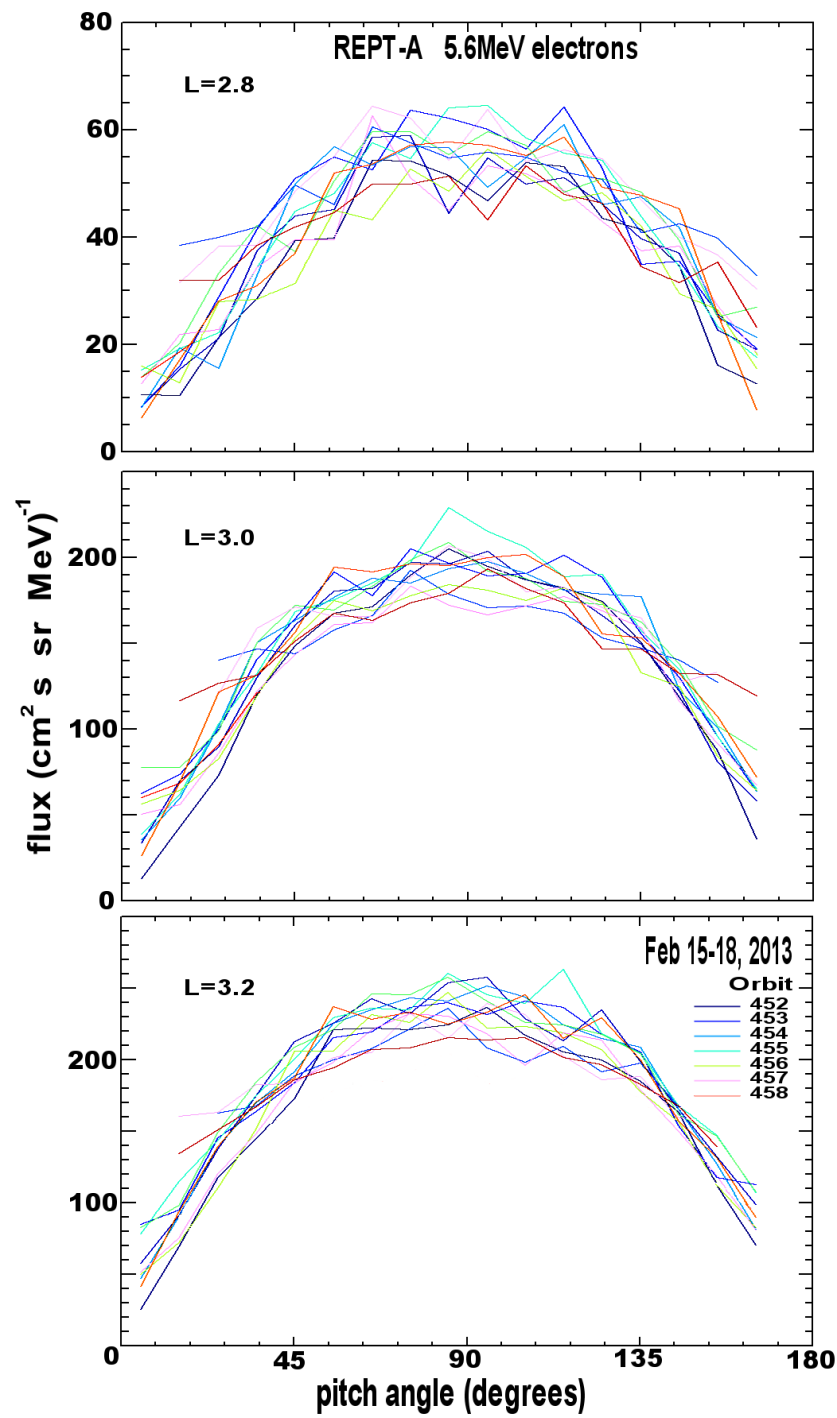


Ultra-Relativistic Electrons: Inner Edge

REPT A & B 7.2 MeV

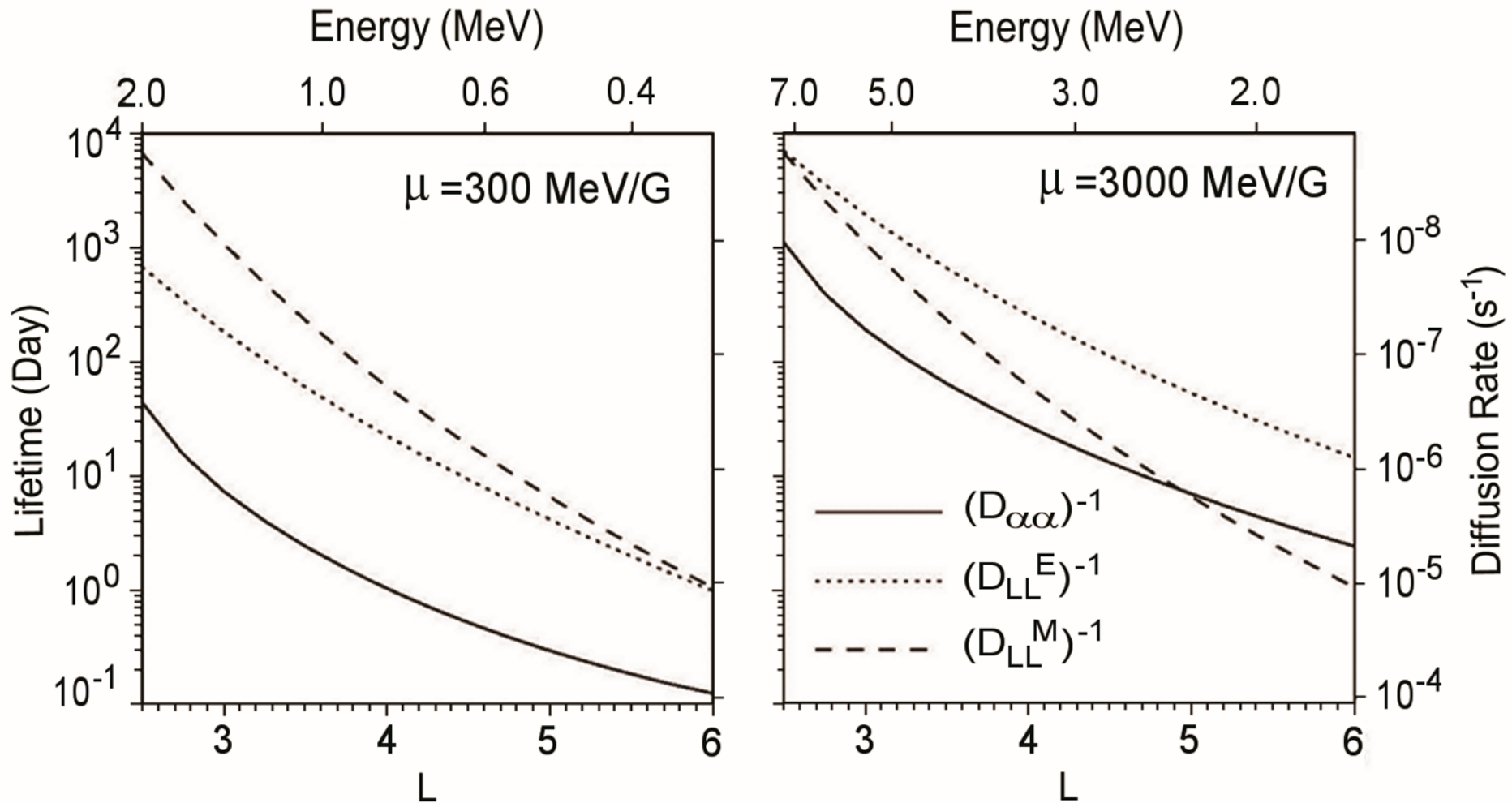




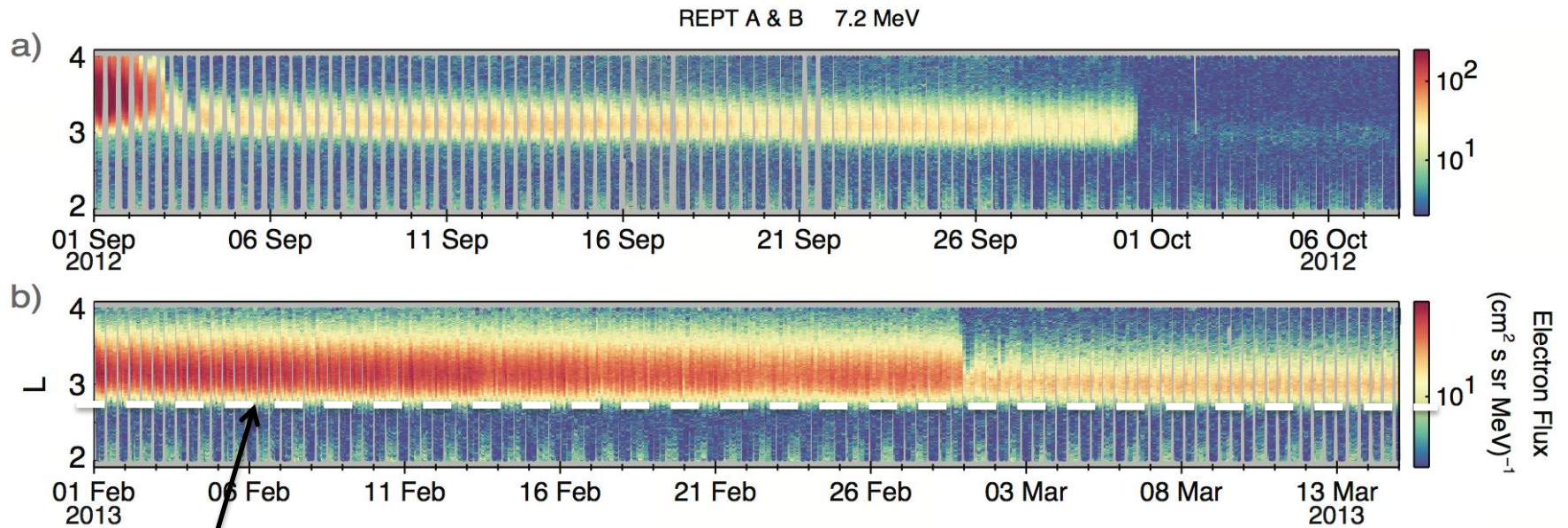


Electron Lifetime Estimates

(Qianli Ma Modeling Results)



An Impenetrable Barrier?



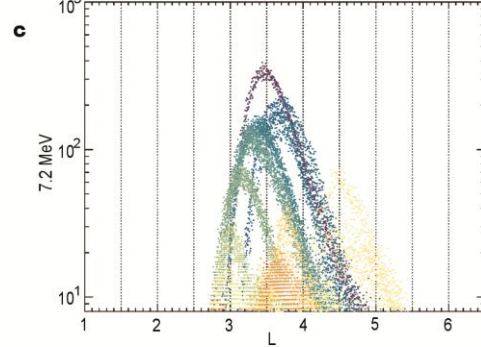
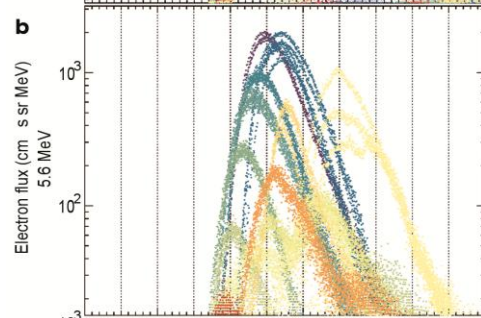
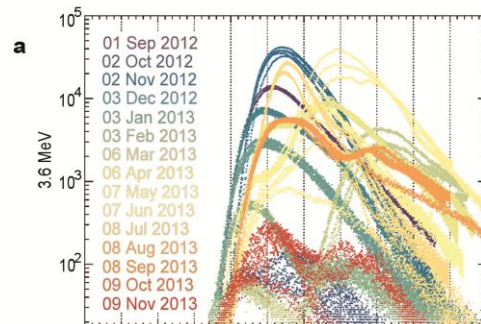
Ultrarelativistic electrons seem never to penetrate inward of $L \sim 2.8$

Conclusions

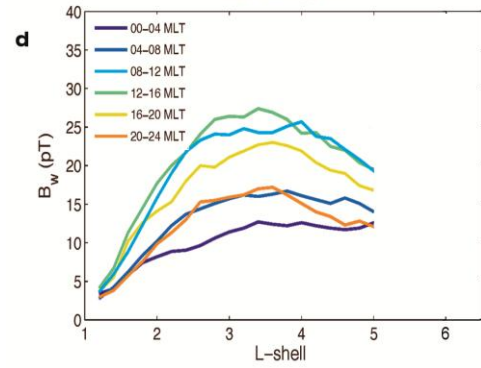
- Results from the Van Allen Probes mission have been rewriting the textbooks about radiation belt structure, acceleration, transport, and loss.
- Excellent energy and pitch angle data reveal distinctive behavior in several electron energy regimes: Highly; Super; and Ultra (> 5 MeV) relativistic.
- REPT data clearly show there is an impenetrable barrier to inward penetration of ultra-relativistic electrons at $L \sim 2.8$ [Baker et al., Nature, 2014].
- The barrier is extremely sharp and is caused neither by geomagnetic field features nor by manmade radio transmissions.
- **A new window has been opened on understanding strong plasma physical gradients in Earth's magnetosphere and this has important significance for remote cosmic systems.**

Thank you.

Questions?



$AL^* > -100$ nT



REPT: The First 19 Months

